On the relative absorption strengths of water vapour in the blue wavelength range


1 Institute of Environmental Physics, University of Heidelberg, Heidelberg, Germany

2 now at: Max Planck Institute for Chemistry, Mainz, Germany

Correspondence to: J. Lampel (johannes.lampel@iup.uni-heidelberg.de)

Received: 19 April 2015 – Published in Atmos. Meas. Tech. Discuss.: 16 June 2015
Revised: 15 September 2015 – Accepted: 29 September 2015 – Published: 15 October 2015

Abstract. In recent updates of the HITRAN water vapour H$_2$O spectroscopic compilation covering the blue spectral region (here: 394–480 nm) significant changes for the absorption bands at 416 and 426 nm were reported. In order to investigate the consistency of the different cross-sections calculated from these compilations, H$_2$O vapour column density ratios for different spectral intervals were retrieved from long-path and multi-axis differential optical absorption spectroscopy (DOAS) measurements. We observed a significant improvement of the DOAS evaluation when using the updated HITRAN water vapour absorption cross-sections for the calculation of the reference spectra. In particular the magnitudes of the residual spectra as well as the fit errors were reduced.

However, we also found that the best match between measurement and model is reached when the absorption cross-section of groups of lines are scaled by factors ranging from 0.5 to 1.9, suggesting that the HITRAN water vapour absorption compilation still needs significant corrections. For this spectral region we present correction factors for HITRAN 2009, HITRAN 2012, HITEMP and BT2 derived from field measurements. Additionally, upper limits for water vapour absorption in the UV-A range from 330 to 390 nm are given.

1 Introduction

Precise knowledge of the atmospheric water vapour absorption not only is crucial to calculate Earth’s radiation budget since water vapour is the largest contributor to the natural greenhouse effect (Myhre et al., 2013) but also needs to be known at high precision when retrieving other atmospheric absorbers from spectroscopic measurements in the UV/VIS wavelength range. If their absorptions overlap, this can lead to errors in measurements of several trace gases. Due to technical limitations and since the absorption cross-sections are small in the region below 500 nm (above 20 000 cm$^{-1}$), the body of available experimental data in this spectral interval is sparse. Numerous absorption lines in the blue spectral region included in absorption line databases like HITRAN 2009 (Rothman et al., 2009), HITRAN 2012 (Rothman et al., 2013) and HITEMP (Rothman et al., 2010) are therefore based on theoretical calculations only. According to Tennyson et al. (2013) and references therein, validation measurements of individual water absorption lines are only available at wavelengths longer than 394.7 nm (25 337 cm$^{-1}$), the body of available experimental data in this spectral interval is sparse. Numerous absorption lines in the blue spectral region included in absorption line databases like HITRAN 2009 (Rothman et al., 2009), HITRAN 2012 (Rothman et al., 2013) and HITEMP (Rothman et al., 2010) are therefore based on theoretical calculations only. According to Tennyson et al. (2013) and references therein, validation measurements of individual water absorption lines are only available at wavelengths longer than 394.7 nm (25 337 cm$^{-1}$), but a large fraction of the absorption lines included in the databases in this region is not validated by measurements.

More recently, water vapour absorption was also observed (Maksyutenko et al., 2012) and modelled (Polyansky et al., 2012) in the UV range below 370 nm. The HITEMP absorption line database for H$_2$O (Rothman et al., 2010) relies to a large extent on the ab initio line list BT2 calculated by Barber et al. (2006). Including the BT2 data increased the number of absorption lines especially in the blue wavelength region drastically: from 410 to 434 nm (W1 + W2, see Table 5) HITRAN 2009 contains 20 absorption lines with in-
tensities above $10^{-27}$ cm molec$^{-1}$ at room temperature, and HITRAN and BT2 contain 382 individual absorption lines. Ratios of absorption line intensities from BT2 and HITRAN are found between 0.5 and 3, because experimental data included in HITRAN 2009 were only combined with the BT2 data when certain criteria were met, as described in Rothman et al. (2010). An overview of the cross-sections from HITRAN 2009, HITEMP and BT2 is shown in Fig. 1. Convoluted to a resolution of 0.5 nm, the differences between the HITRAN 2012 cross-section and the HITEMP cross-section are negligible; therefore they are treated as equivalent here.

The vibrational state of the triatomic asymmetric top water molecule can be described in standard normal mode notation by three quantum numbers: $v_1$ (symmetric stretch), $v_2$ (bend) and $v_3$ (asymmetric stretch). The water vapour absorption shown in Fig. 1 has its structure due to the closeness of the numerous interacting vibrational states. These absorption structures can be named using a convention described in (Tennyson et al., 2001): a polyad is given a label $n\nu$, where $n$ is $v_1 + v_2/2 + v_3$ for even $v_2$ and $n\nu + \delta$ for odd $v_2$.

A widely used approach to measure trace gases in the atmosphere is the technique of differential optical absorption spectroscopy (DOAS) (Platt and Stutz, 2008). Typically, absorption spectra covering 100–200 nm sections in the UV or visible spectral range are recorded by a grating spectrometer at a spectral resolution of 0.3–1 nm. Suitable evaluation of differential absorption structures allows quantitative determination of column densities and concentrations of numerous trace gas species, even if their spectral features overlap. This allows for simultaneous evaluation of several trace gases in the atmosphere at high temporal resolution.

Absorption path lengths for different DOAS setups range from several hundreds of metres to several tens of kilometres in the free troposphere. Optical densities (ODs) can often be resolved down to $1 \times 10^{-4}$ (e.g. Platt and Stutz, 2008; Coburn et al., 2011), which in turn means that water absorption lines can be observed up to an accuracy of $5 \times 10^{-28}$ cm$^2$ molec$^{-1}$ at a resolution of typically 0.5 nm when other absorbers in the respective spectral region are well known and/or their total absorptions are small.

While absorption structures due to known absorbers can be readily disentangled, problems can arise from unknown spectral features. One example – as mentioned above – is the uncertainty of the weak absorption features of water vapour in the blue spectral region.

Frequently the actual water vapour content during the measurements is not of primary interest, but nevertheless water needs to be corrected for when evaluating the absorption spectra. For example, weak water absorptions can be found in the typical evaluation range of iodine monoxide (ca. 414–440 nm). Iodine monoxide (IO) plays a role in atmospheric chemistry and creation of cloud condensation nuclei. Elevated IO levels can be found in coastal areas (e.g. Seitz et al., 2010) as well as on the open ocean (e.g. Read et al., 2008; Großmann et al., 2013), in the free troposphere (e.g. Dix et al., 2013) and in polar regions (e.g. Friess et al., 2010; Schönhardt et al., 2012). For a typical tropospheric water vapour slant column density (SCD) found in moderate climate of $S_{H_2O} = 4 \times 10^{23}$ molec cm$^{-2}$ the resulting optical density is $2.5 \times 10^{-3}$ (when taking the cross-section data from HITEMP) at a spectral resolution of 0.5 nm. Optical densities due to IO in the tropical marine boundary layer are 5 times weaker, i.e. of the order of $4 \times 10^{-4}$; thus to avoid possible errors in the retrieval of IO due to incomplete correction of the water vapour absorption, precise knowledge of the water vapour absorption cross-section is essential. Older versions of HITRAN such as HITRAN 2000 did not even include any absorptions in this spectral range.

Even more important is the correction of water vapour absorptions for the retrieval of glyoxal (e.g. Sinreich et al., 2010; Mahajan et al., 2014, and references therein), which can be evaluated in the spectral range from 432 to 458 nm. For a typical tropospheric water vapour SCD $S_{H_2O} = 4 \times 10^{23}$ molec cm$^{-2}$ the resulting OD is $1.7 \times 10^{-2}$ in this wavelength range at a spectral resolution of 0.5 nm, while the typical glyoxal differential slant column densities (dSCDs) as reported by Sinreich et al. (2010) for the marine boundary layer of $1.5 \times 10^{15}$ molec cm$^{-2}$ correspond to $8 \times 10^{-4}$ and thus is 10 times weaker than the water vapour absorption. Moreover, the main spectral absorption features of water vapour and glyoxal overlap in some parts, potentially introducing cross-sensitivities and/or increasing the measurement error. At the spectral region of maximum absorption of glyoxal the difference between the water vapour absorption cross-sections found in HITRAN 2000 and HITEMP for a typical SCD $S_{H_2O} = 4 \times 10^{23}$ molec cm$^{-2}$ is about $4 \times 10^{-4}$ and amounts
to one-half of the reported glyoxal absorption. Recently the accuracy of glyoxal measurements by CE-DOAS and MAX-DOAS was estimated from chamber measurements by Thalman et al. (2015) and field measurements by Volkamer et al. (2015). For their instruments, water vapour absorption can explain an uncertainty of 5 pptv glyoxal (Volkamer et al., 2015) and less than 15 pptv glyoxal (Thalman et al., 2015).

Water vapour absorption can also have an impact on the spectral retrieval of the oxygen dimer $O_2$ (also called $O_2-O_2$) around 477 nm. Its column densities can be used to constrain radiative transfer simulations in remote sensing application and to obtain height profile information of aerosol extinction and trace gas concentrations (e.g. Frieß et al., 2006).

Water vapour column densities have been measured in the blue wavelength range from satellite (Wagner et al., 2013b; Wang et al., 2014) as well as from ground-based instruments (Wagner et al., 2013a). Here the $7\nu$ absorption at 442 nm from HITRAN 2004 has been used for retrieving the atmospheric water vapour column density. These measurements require precise water vapour absorption cross-sections as well to minimise retrieval errors.

If significant absorptions of water vapour were present below 370 nm, these could have an effect on the spectral evaluation of measurements of e.g. BrO, HCHO, HONO, SO$_2$ and O$_3$. Based on currently available cross-section data, no atmospheric water vapour absorptions in this spectral range have been reported for DOAS measurements so far.

2 Measurement campaigns

The data analysed here were collected during two field campaigns (see Table 1) in which different instruments were used. The DOAS measurements during both campaigns, research cruise SOPRAN M91 off the coast of Peru and HaloCaVe on Cape Verde, aimed to quantify the abundance of reactive trace gases, especially IO and BrO, in the marine boundary layer. HaloCaVe took place parallel to SOPRAN cruise P399 (Bange, 2011) on RV Poseidon in the Mauritanian upwelling region (Hepach et al., 2014). The absorption of IO is overlaid by water vapour absorption. Therefore proper water vapour correction in the spectral retrieval of IO is crucial to obtain reliable measurements of IO. SOPRAN is embedded in the international Surface Ocean – Lower Atmosphere Study (SOLAS) project.

1. During the SOPRAN cruise M91 with the research vessel Meteor multi-axis (MAX)-DOAS data were collected in the Peruvian upwelling region in December 2012 (Bange, 2013).
2. Long-path (LP)-DOAS measurements were analysed for water vapour using data from the intensive campaign HaloCaVe within SOPRAN at the Cape Verde Atmospheric Observatory (CVAO. Carpenter et al., 2010) during summer and fall 2010.

Table 1. Information about the campaigns from which measurement data were used.

<table>
<thead>
<tr>
<th>Name</th>
<th>Type</th>
<th>Location, time</th>
</tr>
</thead>
<tbody>
<tr>
<td>M91</td>
<td>MAX-DOAS</td>
<td>Peru, coastal upwelling, 5° S 82° W–16° S 75° W, 1–25 December 2012</td>
</tr>
<tr>
<td>HaloCaVe</td>
<td>LP-DOAS</td>
<td>CVAO, 16°52’ N 24°52’ W, June–October 2010</td>
</tr>
</tbody>
</table>

The sites of both campaigns were located far away from any strong anthropogenic pollution; thus interferences due to e.g. high NO$_2$ mixing ratios on the data evaluation should be negligible.

3 The DOAS method

The DOAS method (Platt and Stutz, 2008) relies on attenuation of light from suitable light sources by absorbers within the light path according to Lambert–Beer’s law $I(\lambda) = I_0(\lambda) \exp(-\tau(\lambda))$.

The optical density $\tau(\lambda)$ is calculated from a reference spectrum, $I_0(\lambda)$, and a measurement spectrum, $I(\lambda)$, $\tau(\lambda) = -\ln \frac{I(\lambda)}{I_0(\lambda)}$. In order to remove broad-band Mie and Rayleigh extinction, the OD is subdivided into a narrow-band (differential) and a broad-band part, $\tau(\lambda) = \tau_d(\lambda) + \tau_b(\lambda)$. $\tau_d(\lambda)$ is expressed by a sum of the differential parts of possible absorbers with their differential absorption cross-sections $\sigma_{d,i}(\lambda)$ and concentrations $c_i$ of absorber (i.e. trace gas) $i$.

$$\tau_d(\lambda) = L \sum_i c_i \sigma_{d,i}(\lambda)$$ (1)

The column density $S_i = L \times c_i$ of the trace gas $i$ (with the concentration $c_i$) is calculated by a fitting routine, which is applied to data from a wavelength interval with a width of several nanometres to several tens of nanometres. The absorption path $L$ is known for LP-DOAS measurements and can be estimated or calculated from radiative transfer models for MAX-DOAS measurements. The high-resolution literature cross-sections $\sigma_{l,i}$ are convolved with the instrument function $H$ of the respective setup to obtain $\sigma_i = H \otimes \sigma_{l,i}$, the absorption cross-section as it would be determined by the instrument. In an analogous fashion to the optical density, cross-sections can also be subdivided into a broadband and narrow-band (differential) contribution: $\sigma_i = \sigma_{b,i} + \sigma_{d,i}$. The instrument function $H$ is usually measured by observing individual atomic emission lines of mercury, which have a width which is 2 orders of magnitude smaller than the resolution of the instrument (Sansonetti et al., 1996). Corrections to this simple convolution procedure are discussed in Sect. 4.4.
3.1 MAX-DOAS measurements

The method of MAX-DOAS measurements was first described by Höninger and Platt (2002) and uses scattered sunlight collected by a telescope pointing towards the sky at different elevation angles. Each elevation angle has a different sensitivity for absorptions in different heights of the atmosphere. Low elevation angles have a higher sensitivity to absorbers close to the surface, because the additional light path compared to a zenith spectrum is mostly located within the lowermost layers of the atmosphere (Hönninger et al., 2004).

The SCD is defined as the integral over the concentration $\rho$ along the light path $L$ and is hence given in units of molecules cm$^{-2}$. In equations we abbreviate it with $S$.

$$S = \int \rho(s) ds$$

From the MAX-DOAS measurements dSCDs $\Delta S$ can be calculated for each fitted trace gas: a Fraunhofer reference spectrum $I_0(\lambda)$ is chosen from one of the measurement spectra and the dSCD $\Delta S(\alpha) = S(\alpha) - S_{re}t$ is obtained from the DOAS fit for each elevation angle $\alpha$ relative to the Fraunhofer reference. Typically, a zenith spectrum is taken as reference and thus $S_{ref} = S(90^\circ)$. In the measurements reported here, the DOAS fit includes the convolved cross-sections listed in Table 2. In addition, Ring spectra are fitted (see Table 2), which serve to compensate for the effect of rotational Raman scattering. The influence of vibrational Raman scattering (VRS) of $N_2$ and $O_2$ as described in Lampel et al. (2015) on the obtained results listed in Table 5 was not found to be significant. As also described in Lampel et al. (2015), no significant contribution to observed ODs by VRS of liquid water (Vountas et al., 2003; Peters et al., 2014) was found most likely due to turbid water and consequently short light paths under water. VRS of water vapour itself is expected to result in an intensity offset and in differential structures especially at 459 and 463 nm, as it red shifts the solar spectrum by $\approx 3654$ cm$^{-1}$ (Penney and Lapp, 1976; Murphy, 1978; Rizi et al., 2004). However, its expected magnitude ($< 5 \times 10^{-5}$) is below the typical magnitudes of residual spectra. Its spectral signature was not found in these MAX-DOAS measurements.

By choosing references recorded shortly before or after the measurement spectrum, the influence of the instrumental instabilities on the result was minimised and the influence of stratospheric absorbers was largely cancelled out (see e.g. Hönninger and Platt, 2002). In order to avoid measurements in which direct sunlight might have entered the telescope at $90^\circ$ elevation in the southern part of the cruise track, spectra at $40^\circ$ are used as reference spectra.

A description of the instrument operated during the SOPRAN cruise M91 can be found in Großmann et al. (2013). The optical resolution of the instrument during this campaign was $0.45$ nm and it covered a spectral range from 324 to 467 nm.

The spectra used here were recorded for 1 min each at eight elevation angles of $90^\circ$ (zenith), 40, 20, 10, 6, 4, 2 and 1$^\circ$ as long as solar zenith angles (SZAs) were $\leq 85^\circ$. The exposure time was adjusted to have spectra at a typical saturation of 50%.

### Table 2. The cross-sections that were used for the spectral retrieval.

<table>
<thead>
<tr>
<th>Absorber</th>
<th>Source</th>
</tr>
</thead>
<tbody>
<tr>
<td>$NO_2$ 294 K</td>
<td>Vandaele et al. (1998) (MAX-DOAS)</td>
</tr>
<tr>
<td>$NO_2$ 294 K</td>
<td>Vandaele et al. (2002) (LP-DOAS)</td>
</tr>
<tr>
<td>$O_3$ 223 K</td>
<td>Serdyuchenko et al. (2014)</td>
</tr>
<tr>
<td>$O_3$ 293 K</td>
<td>Thalman and Volkamer (2013)</td>
</tr>
<tr>
<td>IO</td>
<td>Spiecz et al. (2005)</td>
</tr>
<tr>
<td>Glyoxal</td>
<td>Volkamer et al. (2005) (for tests only)</td>
</tr>
<tr>
<td>Polynomial third order</td>
<td></td>
</tr>
</tbody>
</table>

MAX-DOAS only:
- Ring: DOASIS, Kraus (2006) and Bussemer (1993)
- Ring $\times \lambda^4$: Wagner et al. (2009)
- Add. polynomial: zeroth order

MAX-DOAS UV only:
- HCHO: Chance and Orphal (2011)
- $NO_2$ 223 K: Vandaele et al. (1998)
- $O_3$ 246 K: Serdyuchenko et al. (2014)

3.2 LP-DOAS Measurements

LP-DOAS instruments are based on an artificial light source, typically a xenon lamp or a light-emitting diode, retro reflectors, a telescope and a spectrometer. The light is sent with a telescope across the measurement distance to a retro reflector, which reflects the light back onto the same telescope. It collects the received light and transfers it to a spectrograph. A sequence of background measurements, light-source spectrum measurements without absorption and actual measurement spectra is used to ensure independence of the measured spectra from external sunlight and instrumental instabilities. The LP-DOAS setup has the advantage that the actual light path is well defined and thus concentrations of molecules can be directly derived.; also, measurements at night are possible.

The optical density $\tau(\lambda)$ is calculated from a background corrected light source spectrum and a background corrected atmospheric spectrum and filtered by a binomial high-pass with 1000 iterations. The convolved and high-pass filtered literature cross-sections listed in Table 2 are then fitted to the corrected OD.

A description of the LP-DOAS instrument can be found in Pößler et al. (2010). The light path used for the measurements...
reported was 12.64 km long, similar to the one in Read et al. (2008). The spectral resolution was 0.5 nm.

To exclude the possibility of interferences with daylight, only spectra at a solar zenith angle of more than 90° (night-time) are reported here.

3.3 Spectral retrieval

The H2O cross-sections were calculated using modelled line widths according to Kuntz (1997) from the respective compilation or line list using an extraction program by Christian Frankenberg (Frankenberg, 2005) with a spectral resolution of 1 pm for an ambient temperature of 298 K and 1013 hPa.1 For the BT2 data set, a cross-section file from the ExoMol project (Tennyson and Yurchenko, 2012) 2 was downloaded; here only Doppler line broadening effects were considered (Hill et al., 2013). Variations of temperature and pressure within the range of atmospheric measurements were found not to have a significant effect on our measurements in the blue wavelength region as the bulk of the absorption by water vapour molecules takes place within the boundary layer.

The choice of cross-section to compensate for absorption of the oxygen dimer O2 did not significantly affect the overall result. We tested cross-sections from Hermans et al. (1999), Thalman and Volkamer (2013) and Greenblatt et al. (1990).

The spectral window of this study for MAX-DOAS measurements was limited at the lower end to 398 nm to avoid a strong influence of the Ring effect caused by rotational Raman scattering (Shefov, 1959; Grainger and Ring, 1962). The upper bound (at 461.5 nm) was chosen due to instrumental limitations.

The water vapour absorption cross-section was divided into six spectral regions, W0–W5, before convolution, as listed in Table 3. All other absorbers (Table 2) were fitted normally. An example fit can be seen in Fig. 2.

In the literature, significant absorption structures due to glyoxal in the eastern Pacific region were reported by Sinreich et al. (2010). However, during our cruise no absorption due to glyoxal was found to exceed our detection limit of 2σ = 5 × 1014 molec cm−2 glyoxal at low elevation angles of 1–3° relative to 40°. This detection limit was determined after analysing the spectral data in a fit window from 432 to 460 nm. The upper limit was independent of the choice of literature O4 absorption cross-section and independent of the choice of water vapour cross-section (tested for HITRAN2009, HITEMP and BT2). Furthermore, it was not exceeded when including a correction spectrum for VRS of liquid water, liquid water absorption, VRS of air or a correction spectrum to account the changes of the effective water vapour absorption band shape introduced by radiative transfer effects in strong water vapour absorption lines around 442 nm. The result remained the same for different fit windows.

To avoid problems in situations with low light intensity, we used only data in which the signal was sufficiently high to provide a RMS noise of the residual spectrum below 4 × 10−4.

3.3.1 Radiative transfer effects (MAX-DOAS)

The light path of the LP-DOAS measurement is well defined and constant. Thus measurements of the different W0–W5 column densities can be directly used. However, the effective light path of MAX-DOAS measurements is not known and

Table 3. Integrated absorption in [10−27 nm cm2] over each of the wavelength intervals W0–W5 for different sources of cross-section data. Not only are W3 variations (reference measurements, bold face) between the different compilations seen for the largest absorption structure but relative integrated absorption values also vary. The last row shows the maximum optical density for a water vapour column density (CD) S = 4 × 1023 molec cm−2 within each wavelength interval at a spectral resolution of 0.5 nm.

<table>
<thead>
<tr>
<th>Dominating polyad</th>
<th>8ν</th>
<th>7ν + δ</th>
<th>7ν</th>
<th>6ν + δ</th>
</tr>
</thead>
<tbody>
<tr>
<td>Name</td>
<td>W0</td>
<td>W1</td>
<td>W2</td>
<td>W3</td>
</tr>
<tr>
<td>Start of interval</td>
<td>(nm)</td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>HITRAN 2000</td>
<td>0.00</td>
<td>0.00</td>
<td>0.00</td>
<td>69.02</td>
</tr>
<tr>
<td>HITRAN 2004</td>
<td>13.62</td>
<td>3.11</td>
<td>0.89</td>
<td>96.75</td>
</tr>
<tr>
<td>HITRAN 2009</td>
<td>13.71</td>
<td>3.13</td>
<td>0.90</td>
<td>97.07</td>
</tr>
<tr>
<td>HITEMP</td>
<td>21.01</td>
<td>15.73</td>
<td>4.01</td>
<td>106.90</td>
</tr>
<tr>
<td>BT2</td>
<td>26.05</td>
<td>23.84</td>
<td>7.86</td>
<td>116.50</td>
</tr>
</tbody>
</table>

Source of cross-section data | integrated cross-section
| HITRAN 2000 | [10−27 nm cm2] | 0.00 | 0.00 | 0.00 | 69.02 | 0.00 | 31.03 |
| HITRAN 2004 | 13.62 | 3.11 | 0.89 | 96.75 | 0.87 | 42.25 |
| HITRAN 2009 | 13.71 | 3.13 | 0.90 | 97.07 | 0.88 | 42.46 |
| HITEMP      | 21.01 | 15.73| 4.01 | 106.90| 4.50 | 51.44 |
| BT2         | 26.05 | 23.84| 7.86 | 116.50| 8.46 | 62.67 |

OD HITEMP for CD = 4 × 1023 molec cm−2 [10−4] | 36 | 27 | 6 | 165 | 4.5 | 62 |
The overall water vapour column density calculated from the absorption bands, correlation plots with respect to W3 were performed as shown exemplary in Fig. 3. From all fits with separate column densities for the absorption bands, correlation plots with respect to W3 were performed as shown exemplary in Fig. 3. The results are listed in Table 5. The overall water vapour column density calculated from the data shown above is \(\text{dSCD} = 5.31 \times 10^{23} \text{molec cm}^{-2}\).

Figure 2. Fit residuals (top panel) and water vapour OD (bottom) for a MAX-DOAS observation using the HITEMP cross-section. A residual from an individual fit using the original HITEMP cross-section and a fit using separate column densities for each of the absorption structures in the windows W0–W4 from HITEMP is shown in the top panel. From all fits with separate column densities for the absorption bands, correlation plots with respect to W3 were performed as shown exemplary in Fig. 3. The results are listed in Table 5. The overall water vapour column density calculated from the data shown above is \(\text{dSCD} = 5.31 \times 10^{23} \text{molec cm}^{-2}\).

4 Results

The water vapour dSCDs derived in different spectral regions of the same measured spectra were compared to each other. In principle the same water vapour dSCD values are expected, but significant differences were found.

<table>
<thead>
<tr>
<th>Wavelength (nm)</th>
<th>W0</th>
<th>W1</th>
<th>W2</th>
<th>W3</th>
<th>W4</th>
<th>W5</th>
</tr>
</thead>
<tbody>
<tr>
<td>SD</td>
<td>0.02</td>
<td>0.02</td>
<td>0.01</td>
<td>–</td>
<td>0.02</td>
<td>0.02</td>
</tr>
<tr>
<td>McArtim</td>
<td>0.93</td>
<td>0.96</td>
<td>0.98</td>
<td>1.00</td>
<td>1.02</td>
<td>1.04</td>
</tr>
</tbody>
</table>

Table 4. Corrections of the effective light path according to radiative transfer modelling for the MAX-DOAS measurements. The reference measurements, W3, are in bold.

The results were averaged over a range of SZAs of 10–80° representing the used MAX-DOAS data and over all encountered relative solar azimuth angles. The standard deviation listed in Table 4 reflects the variations in simulated \(\text{H}_2\text{O} \) dSCDs due to varying observation geometries.

The magnitude of the resulting water vapour dSCDs obtained from the model agreed with the MAX-DOAS observations. McArtim allows calculating the wavelength dependence of the simulated water vapour dSCDs. The scatter in the correlations for the water absorption bands (Fig. 3) will then already include the scatter caused by different measuring geometries, which means that the observed differences in relative strengths of the absorptions especially for W1 and W2 are significant. However, the correction obtained from radiative transfer calculations (shown in Table 4) needs to be applied to MAX-DOAS observations to match the LP-DOAS results. The correction of the wavelength dependence reduces the discrepancy of MAX-DOAS and LP-DOAS measurements e.g. for W0 when using HITEMP and BT2. The wavelength-corrected results are listed in Table 5.

It is a well-known effect (see e.g. Platt et al., 1997) that strong absorbers influence the light path length in the atmosphere and thus the air mass factor. The change of the air mass factor within a water vapour absorption band due to the highly resolved water vapour optical density in the region <480 nm was found to be insignificant. Deviations of less than 2 % on the values listed in Table 5 were observed.

However, if the focus is on weak absorbers being overlayed by strong water vapour absorptions, a correction for the change of air mass factors by strong absorption is necessary to avoid water vapour OD-correlated structures in the residual spectra and effects on the retrieval of other absorbers.

J. Lampel et al.: Relative absorption strengths of water vapour absorption
Each cross-section in W0–W5 for individual measurements are summarised in Table 6. In order to illustrate the differences, one of the data comparison plots with the relative absorption of W1 and W3 for HITRAN is shown in Fig. 3.

The weights for the linear regression were iteratively calculated according to Neri et al. (1989). A comparison with other methods for error-weighted linear regressions can be found in Cantrell (2008).

The water vapour absorption at $\approx 444$ nm was used as a reference for the other absorption bands, because it is the strongest absorption in this spectral region (see Table 3). The mathematical error from the regression is small compared to possible systematic errors, which might arise from the DOAS fit itself: its relative size is estimated by using twice the fit error given in Table 6 following the argumentation of Stutz and Platt (1996). These errors are then used to calculate the weighted mean deduced from MAX-DOAS and LP-DOAS observations for each of the absorption bands given in Table 5, when their respective relative fit error is smaller than 25 %.

In order to exclude an impact of glyoxal absorption on the evaluation of water vapour absorptions, the analysis procedure was repeated including glyoxal. The result is comparable to the result without glyoxal absorption, only the amplitude of the weak water absorption around 426 nm was slightly reduced by the additional degree of freedom introduced to the spectral retrieval.

For the LP-DOAS measurements an upper limit of 30 ppt glyoxal (see also Mahajan et al., 2014) was determined and its cross-section was also not included in the final analysis.

Other MAX-DOAS observations in the same region showed larger glyoxal dSCDs (Sinreich et al., 2010), which was also detected at various other campaigns in the marine boundary layer as listed in Mahajan et al. (2014) and Volka-

Table 5. Measured relative line strengths for the different cross-sections with respect to the absorption at W3 (see Fig. 1 and Table 3), which is the reference column in bold face. Errors obtained from the linear regression are shown for the last digits in brackets. The relative DOAS fit errors are listed in Table 6. Results with typical DOAS fit errors of more than 25 % of the measured values were put in square brackets. MAX-DOAS values are corrected by the results of radiative transfer modelling listed in Table 4.

<table>
<thead>
<tr>
<th>Name</th>
<th>$8\nu$</th>
<th>$7\nu + \delta$</th>
<th>$7\nu$</th>
<th>$6\nu + \delta$</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>W0</td>
<td>W1</td>
<td>W2</td>
<td>W3</td>
</tr>
<tr>
<td>Start of interval (nm)</td>
<td>394.0</td>
<td>410.0</td>
<td>423.5</td>
<td>434.0</td>
</tr>
<tr>
<td>End of interval (nm)</td>
<td>410.0</td>
<td>423.5</td>
<td>434.0</td>
<td></td>
</tr>
</tbody>
</table>

Source of cross-section data

<table>
<thead>
<tr>
<th>Source of cross-section data</th>
<th>DOAS</th>
<th></th>
<th></th>
<th></th>
<th></th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>HITRAN 2009</td>
<td>MAX</td>
<td>1.0875(5)</td>
<td>1.9497(14)</td>
<td>[1.6862(46)]</td>
<td>1</td>
<td>[0.3115(51)]</td>
</tr>
<tr>
<td></td>
<td>LP</td>
<td>1.02(2)</td>
<td>0.99(6)</td>
<td>[1.6(2)]</td>
<td>1</td>
<td>[0.7(3)]</td>
</tr>
<tr>
<td>HITEMP</td>
<td>MAX</td>
<td>1.0201(4)</td>
<td>0.6578(4)</td>
<td>0.769(2)</td>
<td>1</td>
<td>0.360(2)</td>
</tr>
<tr>
<td></td>
<td>LP</td>
<td>1.17(2)</td>
<td>0.42(2)</td>
<td>[0.91(8)]</td>
<td>1</td>
<td>1.03(1)</td>
</tr>
<tr>
<td>HITEMP (with glyoxal)</td>
<td>MAX</td>
<td>1.0182(4)</td>
<td>0.6534(4)</td>
<td>0.744(2)</td>
<td>1</td>
<td>0.359(2)</td>
</tr>
<tr>
<td></td>
<td>LP</td>
<td>1.0108(4)</td>
<td>0.5183(3)</td>
<td>0.546(1)</td>
<td>1</td>
<td>0.395(1)</td>
</tr>
<tr>
<td>BT2</td>
<td>MAX</td>
<td>1.0108(4)</td>
<td>0.5183(3)</td>
<td>0.546(1)</td>
<td>1</td>
<td>0.395(1)</td>
</tr>
<tr>
<td></td>
<td>LP</td>
<td>1.19(2)</td>
<td>0.37(2)</td>
<td>[0.74(8)]</td>
<td>1</td>
<td>1.01(1)</td>
</tr>
</tbody>
</table>

Values of the retrieved water vapour SCDs in different spectral intervals were compared to each other. Relative to the differential SCD of W3 $S_{W3}$, an error-weighted linear regression was done to obtain $a_x$ in $f(S_x) = a_x \times S_{W3}$. $a_x$ is the relative absorption strength of the interval $x$ relative to W3. The results are shown in Table 5. Typical DOAS fit errors for each cross-section in W0–W5 for individual measurements are summarised in Table 6. In order to illustrate the differences, one of the data comparison plots with the relative absorption of W1 and W3 for HITEMP is shown in Fig. 3.

The mathematical error from the regression is small compared to possible systematic errors, which might arise from the DOAS fit itself: its relative size is estimated by using twice the fit error given in Table 6 following the argumentation of Stutz and Platt (1996). These errors are then used to calculate the weighted mean deduced from MAX-DOAS and LP-DOAS observations for each of the absorption bands given in Table 5, when their respective relative fit error is smaller than 25 %.

In order to exclude an impact of glyoxal absorption on the evaluation of water vapour absorptions, the analysis procedure was repeated including glyoxal. The result is comparable to the result without glyoxal absorption, only the amplitude of the weak water absorption around 426 nm was slightly reduced by the additional degree of freedom introduced to the spectral retrieval.

For the LP-DOAS measurements an upper limit of 30 ppt glyoxal (see also Mahajan et al., 2014) was determined and its cross-section was also not included in the final analysis.

Other MAX-DOAS observations in the same region showed larger glyoxal dSCDs (Sinreich et al., 2010), which was also detected at various other campaigns in the marine boundary layer as listed in Mahajan et al. (2014) and Volka-
The detection limit found here is at the lower range of these observations. As we performed various sensitivity studies for this data set, which could not explain this difference, we assume that this difference could be due to natural variability. The bulk of the measurements published in Sinreich et al. (2010) was not as close to the coast as M91. The measurements by Volkamer et al. (2015) were done in a different region, even north of the equator, which additionally shows typically larger sea surface temperatures (SSTs) than the data presented here (SST during M91: 14–22 °C; Fuhlbrügge et al., 2015).

The impact on IO dSCDs of the rescaled water vapour absorption is small, as long as the main water vapour absorption feature W3 is not included in the fit interval, which is typically the case as (e.g. Großmann et al., 2013; Prados-Roman et al., 2015; Volkamer et al., 2015). The impact on the IO dSCD is found to be smaller than 1 × 10^{12} molec cm\(^{-2}\) for a fit interval from 418 to 438 nm.

### 4.1 Different separation approaches

Deriving water vapour dSCDs in different wavelength windows is an approach which yields direct information about the relative absorption strength of different absorption bands. The result can easily be applied to DOAS measurements.

However, as the different polyads described by different \( n = v_1 + v_2/2 + v_3 \) for the vibrational states \((v_1, v_2, v_3)\) of the water molecule overlap in the blue wavelength region, a separation by polyad could yield further information.

In the upper panel of Fig. 4, the HITEMP line list was separated according to its polyads and convolved to a representative spectral resolution of 0.5 nm. Absorption lines which are not assigned to vibrational quantum numbers in HITEMP and can therefore not be assigned to one of the polyads are labelled “NA” in this plot. These absorption lines alone amount to an OD of several 10\(^{-4}\) for a dSCD of 4 × 10\(^{23}\) molec cm\(^{-2}\). Then the same separation was done for different values of \( v_2 \).

Using the same spectral retrieval for MAX-DOAS measurements as described in Sect. 3.3 but using now the polyads \( 6\nu + \delta, 7\nu, 7\nu + \delta \) and the list of not-assigned (NA) absorption lines instead of W0–W5 did not show a significant improvement of the residual of the fit (mean RMS with W0–W5: \( (1.89 \pm 0.01) \times 10^{-4} \); mean RMS with polyads: \((1.90 \pm 0.01) \times 10^{-4}\)). This could have been caused by the NA absorption lines, which might scale differently. Or the separation of the absorption bands according to \( n \) is not representing the underlying problem. A further indication, that this is indeed the case is the fact that the scaling of W2 is typically not close to unity, but here it belongs to \( 7\nu \), which is by definition “correct”. The absorption lines of W2 belong formally to the \( 7\nu \) polyad, but as their bend modes are mostly \( v_2 = 0, 2 \), this difference might in fact drive the observed differences. The W1 absorptions, which are overestimated in HITEMP, consist mostly of \( v_2 = 1, 2 \), as also shown in Fig. 4. As the \( v_2 = 1 \) absorption lines seem to be of correct size (as indicated by W5), the reason for the observed discrepancies might be connected to higher bend states with \( v_2 \geq 2 \).

Table 7 shows that the scaling factor for the polyads is close to unity except for \( 7\nu + \delta \). As different scaling factors for the \( 7\nu \) and \( 7\nu + \delta \) polyads and “NA” will lead to a different shape of the absorption band W1 around 416 nm, this could explain the remaining residual structures shown in Fig. 2 when rescaling the whole absorption band W1. However, if the structure in \( 7\nu \) at 416 nm is not correct, this could be partly compensated by \( 7\nu + \delta \) polyad, which then appears to have an overestimated absorption cross-section, as it is the case here.

In the lower panel in Fig. 4, the HITEMP line list is separated according to bend values \( v_2 \). Here it is obvious why the absorption W2 was labelled \( 7\nu \), as here \( v_2 = 2 \). In the same way as the polyads were evaluated, the result from this

### Table 6. Typical DOAS fit errors in fitting windows W0–W5 at a water vapour dSCD in W3 (bold) of 4 × 10\(^{23}\) molec cm\(^{-2}\) for an individual spectrum integrated over 60 s (MAX-DOAS) and 280 s (averaged, LP-DOAS). Values are given in percent and are corrected by the relative sizes given in Table 5. For all weak absorption bands W1, W2 and W4 a reduction of fit errors is observed from HITRAN 2009 to HITEMP or BT2.

<table>
<thead>
<tr>
<th>(%)</th>
<th>Start of interval (nm)</th>
<th>W0</th>
<th>W1</th>
<th>W2</th>
<th>W3</th>
<th>W4</th>
<th>W5</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>End of interval (nm)</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>Source of cross-section data</td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>HITRAN 2009</td>
<td>MAX-DOAS</td>
<td>2.9</td>
<td>5.1</td>
<td>20</td>
<td>0.63</td>
<td>130</td>
<td></td>
</tr>
<tr>
<td></td>
<td>LP-DOAS</td>
<td>5.2</td>
<td>21</td>
<td>41</td>
<td>1.04</td>
<td>132</td>
<td>3.6</td>
</tr>
<tr>
<td>HITEMP</td>
<td>MAX-DOAS</td>
<td>2.6</td>
<td>3.9</td>
<td>13</td>
<td>0.59</td>
<td>42</td>
<td>–</td>
</tr>
<tr>
<td></td>
<td>LP-DOAS</td>
<td>4.7</td>
<td>16</td>
<td>28</td>
<td>0.98</td>
<td>32</td>
<td>3.4</td>
</tr>
<tr>
<td>BT2</td>
<td>MAX-DOAS</td>
<td>3.7</td>
<td>5.4</td>
<td>17</td>
<td>0.85</td>
<td>33</td>
<td>–</td>
</tr>
<tr>
<td></td>
<td>LP-DOAS</td>
<td>4.6</td>
<td>16</td>
<td>28</td>
<td>0.94</td>
<td>34</td>
<td>3.2</td>
</tr>
</tbody>
</table>
Table 7. Relative absorption strengths as listed in Table 5 for the separated H$_2$O absorption cross-section according to the respective polyad shown in Fig. 4. Below the same procedure for the HITeMP line list separated by eigenvalues of the bend mode $v_2$. As in Table 6 the fit error is given in percent at a dSCD of $4 \times 10^{23}$ molec cm$^{-2}$ of the $7v$ and $v = 0$ absorptions respectively. As the wavelength ranges are not as well defined as before, no corrections by the results of radiative transfer modelling listed in Table 4 were done.

<table>
<thead>
<tr>
<th>Polyad</th>
<th>$8v$</th>
<th>$7v + \delta$</th>
<th>$7v$</th>
<th>$6v + \delta$</th>
<th>“NA”</th>
</tr>
</thead>
<tbody>
<tr>
<td>HITEMP</td>
<td>0.977(1)</td>
<td>0.283(1)</td>
<td>1</td>
<td>0.791(2)</td>
<td>1.045(2)</td>
</tr>
<tr>
<td>Fit error (%)</td>
<td>5</td>
<td>19</td>
<td>1</td>
<td>8</td>
<td>10</td>
</tr>
<tr>
<td>$v_2$</td>
<td>0</td>
<td>1</td>
<td>2</td>
<td>3</td>
<td>4</td>
</tr>
<tr>
<td>HITEMP</td>
<td>1</td>
<td>0.583(2)</td>
<td>0.427(3)</td>
<td>0.306(4)</td>
<td>$[0.675(2)]$</td>
</tr>
<tr>
<td>Fit error (%)</td>
<td>0.6</td>
<td>6</td>
<td>12</td>
<td>23</td>
<td>42</td>
</tr>
</tbody>
</table>

Figure 4. The HITeMP line list for water vapour separated according to polyads and bend mode: “NA” summarises all absorption lines which are listed with the vibrational quantum numbers $(v_1, v_2, v_3) = (-2, -2, -2)$ in HITeMP. The dotted horizontal line marks an OD of $1 \times 10^{-4}$ for a dSCD of $4 \times 10^{23}$ molec cm$^{-2}$.

separation was also applied to the MAX-DOAS observations in order to see if this separation agrees better with observations. The results are listed in the lower part of Table 7, and a significant improvement of the RMS was not observed. The dominating $v_2 = 0 - 4$ values were separated and the remaining lines assigned to $v_2 = 5 - 9$ were combined with the NA absorption lines as also shown in Fig. 4.

The resulting scaling factors decrease with increasing values $v_2$ for the bend mode, which suggests that the observed discrepancies of relative absorption strength are caused mainly by higher values of $v_2$. The result for $v_2 = 4$ is close to the limit of detection.

4.2 Comparison of LP-DOAS data with data from meteorological stations

In order determine whether the assumed water vapour absorption cross-sections actually give the correct water vapour concentration, the main absorption W3 found in LP-DOAS data was compared to water vapour mixing ratios derived from meteorological parameters measured at the CVAO.

The meteorological station provides temperature, pressure and humidity data (Carpenter et al., 2010). This information was used to calculate the water vapour mixing ratio using the Magnus formula and to compare the result with LP-DOAS
data. Taking the measurement error given for the meteorological station, the stated error in the temperature measurements of 0.3 °C results in 2% uncertainty in the relative humidity. Pressure uncertainties cancel out, since the same pressure measurements were used for the conversion of column densities from LP-DOAS to mixing ratios as well as in the Magnus formula. An error of 5% in relative humidity directly translates in an 5% error for the mixing ratio. This means that the absolute differences of the cross-sections shown in Table 3 cannot be absolutely validated with sufficient precision, even though the water vapour mixing ratios ranged from 2.0 to 3.4% and meteorological station data and LP-DOAS water vapour data correlated with a Pearson’s $R = 0.93$ as shown in Fig. 5. Slope and offset values are listed in Table 8. The LP-DOAS results based on the HITEMP cross-section were on average 7% lower than the values inferred from the meteorological station data and thus closer to the absolute magnitude of the W3 absorption in HITRAN 2009. Retrieving tropospheric water vapour profiles from the MAX-DOAS measurements introduced even larger deviations probably due to uncertainties in retrieving the required aerosol profiles.

4.3 Absorption of water vapour in the UV wavelength range

As mentioned above, water vapour absorptions are observed not only in the spectral region above 390 nm but also down to wavelengths approaching the dissociative continuum starting below 243 nm (Maksyutenko et al., 2012). Also ab initio calculations (Polyansky et al., 2012) for this spectral region were developed.

BT2 and HITEMP include absorption features of water vapour in the wavelength range below 390 nm, as shown in Fig. 1. Other trace gases such as HCHO, BrO and HONO are retrieved in this wavelength region. However, to date, water vapour absorptions are not included in DOAS retrieval procedures. Therefore the detection of water vapour absorption below 390 nm and thus the possibility to improve trace gas retrievals are of great interest for DOAS measurements.

For the DOAS analysis in the UV range, IO was removed from the fit while the cross-sections listed in Table 2 were added.

During the MAX-DOAS measurements (M91) systematic residual structures in the spectral region below 370 nm were observed. Their amplitude of up to $6 \times 10^{-4}$ increased with decreasing elevation angle just like the O$_4$ and/or water vapour dSCDs (at 442 nm) and also the residual spectral structure showed narrow differential absorption features. This could indicate a tropospheric absorber in the spectral region below 370 nm – which is currently not considered in typical DOAS retrievals – such as water vapour. Due to the strong correlation of O$_4$ and water vapour dSCDs for the MAX-DOAS measurements during the M91 campaign and insufficient detection limits for the LP-DOAS measurements, these residual structures could not be unambiguously attributed to either O$_4$ or water vapour absorption. The variation of water vapour mixing ratios along the cruise track of M91 was not large enough to separate both possible contributions.

The water vapour absorption band around 377 nm could not be identified by our MAX-DOAS measurements (M91) so far and is, judging from the residual optical depth from the fit, presumably smaller than $4 \times 10^{-27}$ cm$^2$ at a spectral resolution of 0.45 nm. This is in agreement with the published values of the cross-sections shown in Fig. 1.

The specified cross-section of the absorption band at 362 nm is about twice as large in BT2 as in HITEMP, which is due to the line cutoff present in the HITEMP database. The absorption lines around 362 nm are not based on measurements but on the calculated BT2 line list, which was the starting point for the HITEMP database (Rothman et al., 2010). The effect of the cutoff in HITEMP can amount to an OD of $4 \times 10^{-4}$ for a SCD $S_{H_2O} = 4 \times 10^{23}$ molec cm$^{-2}$ as seen from the difference of the convolved absorptions listed in BT2 and HITEMP around 362 nm. Below 394 nm no laboratory measurements of individual water vapour absorption bands are available, as listed in Tennyson et al. (2013). Thus the absorption lines listed in HITEMP below 394 nm originate from BT2.

Furthermore, the retrieval of this band from atmospheric spectra would be difficult due to uncertainties of the overlaying absorption of the collision-induced absorption by the oxygen dimer O$_4$. Independent of the employed literature O$_4$ cross-section in our DOAS evaluation, the suspected water vapour absorption appears to be present in our measurement data, but the instability of the fit with respect to the used fit in-
Table 8. Comparison of LP-DOAS data with data from the meteorological station at the CVAO: the measured mixing ratio \( m \) is compared to the calculated mixing ratio \( w \) from the data of the weather station. \( f(w) = a \cdot w + c \) and \( f_0(w) = a_0 \cdot w \) are fitted measurement error-weighted to the observations \( m \). The correlation coefficient or Pearson’s \( R \) is constant for all cases. The best agreement in overall absorption size in the W3 interval is observed for HITRAN 2009. The uncertainty of the meteorological measurements is estimated to result in a relative uncertainty in water vapour mixing ratio of \( 7\% \). The relative observed slopes \( a \) and \( a_0 \) agree with expectations from Table 3.

<table>
<thead>
<tr>
<th>Source of cross-section data</th>
<th>( a )</th>
<th>( c (%) )</th>
<th>( a_0 )</th>
<th>( R )</th>
</tr>
</thead>
<tbody>
<tr>
<td>HITRAN 2009</td>
<td>1.02 ( \pm ) 0.03</td>
<td>0.45 ( \pm ) 0.80</td>
<td>1.004 ( \pm ) 0.003</td>
<td>0.93</td>
</tr>
<tr>
<td>HITEMP</td>
<td>0.93 ( \pm ) 0.03</td>
<td>0.40 ( \pm ) 0.70</td>
<td>0.918 ( \pm ) 0.002</td>
<td>0.93</td>
</tr>
<tr>
<td>BT2</td>
<td>0.90 ( \pm ) 0.03</td>
<td>0.41 ( \pm ) 0.69</td>
<td>0.882 ( \pm ) 0.002</td>
<td>0.93</td>
</tr>
</tbody>
</table>

terval and large residual structures close to the possible water absorption indicate that there is still a significant mismatch between measurement and modelled absorption lines.

Du et al. (2013) reported measurements of the water vapour absorption cross-section by ring-down spectroscopy of pure water vapour in the 290–350 nm region sampling the absorption cross-section in steps of 5 nm. They report that their measurements are in agreement with previous determinations of the absorption at 442.73 nm. The reported cross-section values of \( \sigma_{\text{max}} = 2.94 \times 10^{-24} \text{cm}^2 \text{molec}^{-1} \) at 330 nm exceed the maximum absorption of the BT2 line list in the spectral region from 330 to 350 nm by 2 orders of magnitude. This would lead to an OD of \( \tau = S \times \sigma_{\text{max}} = 1.2 \) for MAX-DOAS measurements with a SCD \( S_{\text{H}_2\text{O}} = 4 \times 10^{23} \text{molec cm}^{-2} \) (a typical MAX-DOAS column density for mid-latitude summer conditions) under a telescope elevation angle of \( 3^\circ \), which is clearly in disagreement with our observations. Also measurements of other tropospheric trace gases such as HCHO (336.5–359 nm) (Pinardi et al., 2013), BrO (330.6–352.75 nm) and SO\(_2\) (314.8–326.8 nm) (Lübcke et al., 2014), which would have also been affected, did not show any unknown differential absorption of this size but did yield residual spectra with at least 2 orders of magnitude smaller amplitudes.

For MAX-DOAS measurements in the Peruvian upwelling (M91) the magnitude of the fit residual in the region from 332 to 370 nm was below \( 6 \times 10^{-4} \) peak to peak for a water vapour dSCD retrieved in the blue wavelength range of \( 4 \times 10^{23} \text{molec cm}^{-2} \), yielding an upper limit on the differential cross-section of water vapour of \( 3 \times 10^{-27} \text{cm}^2 \text{molec}^{-1} \) at a resolution of 0.45 nm under atmospheric conditions, i.e. 3 orders of magnitude smaller than the figure reported by Du et al. (2013) at 330 nm.

If the shape of the water vapour cross-section presented in Du et al. (2013) could be represented by the DOAS polynomial in the respective wavelength range, it would not be detected. However, the comparably small absorption at 345 nm would have already resulted in a significant absorption structure with a differential absorption structure size of \( 6 \times 10^{-2} \) which could not have been compensated by the polynomial. Additionally, water vapour does not show an absorption cross-section which can be represented by a polynomial in other spectral regions in the visible spectral range. We therefore conclude that the cross-section values reported in Du et al. (2013) must be considerably too high, judging from UV MAX-DOAS measurements under atmospheric conditions. Alternatively the measurements of Du et al. (2013) may represent only individual absorption lines at each of the wavelengths of the reported magnitude, while in the spectral region between those measurements no cross-section data are available. Therefore the conclusion by Du et al. (2013) that the total impact of water vapour absorptions in middle-latitudes on the radiative flux at the ground level is be comparable to ozone between 330 and 350 nm cannot be supported.

At a spectral resolution of 0.5 nm, the OD attributed to water vapour between 350 and 370 nm is according to HITEMP \( 5.5 \times 10^{-4} \) (BT2: \( 9.5 \times 10^{-4} \)) for a typical H\(_2\)O–dSCD of \( 4 \times 10^{23} \text{molec cm}^{-2} \). Under similar measurement conditions (dSCD\(_{\text{O}_3} = 4 \times 10^{43} \text{molec}^2 \text{cm}^{-2} \)) the OD due to \( \text{O}_3 \) is \( 2 \times 10^{-2} \), i.e. by a factor of 20–40 larger. The OD according to HITEMP/BT2 from 330 to 350 nm is comparable to the optical density of 1 ppt BrO along a light path of 10 km and could thus be crucial for DOAS measurements of BrO in the remote marine boundary layer (compare e.g. Volkamer et al., 2015). Furthermore these absorptions could contribute to the observed problems during the retrieval of tropospheric HCHO as described by Pinardi et al. (2013).

To estimate the overall influence of water vapour on the retrieval of \( \text{O}_3 \), BrO, HCHO and HONO further dedicated laboratory measurements of water vapour in this spectral region are needed.

### 4.4 Error sources

The residual spectra obtained from the DOAS evaluation procedure of the measurements were usually not dominated by photon shot noise but showed recurrent residual structures as show in Fig. 2. Since the influence of absorbers other than H\(_2\)O is probably negligible, these residual structures from instrumental instabilities, saturation and \( I_0 \)-effect (for MAX-DOAS) should be the dominating sources of interference. Their influence on the absolute magnitude of the absorption bands is discussed in the following subsections, the influence of radiative transfer effects is discussed in Sect. 3.3.1.
4.4.1 Other absorbers

For MAX-DOAS measurements the OD of absorbers other than water vapour was kept low by using a Fraunhofer reference $I_0(\lambda)$ recorded within the same sequence of elevation angles. Using this approach the solar zenith angle only changed slightly between the measurements and most of the stratospheric absorption by ozone and NO$_2$ cancel out and do not affect the evaluation. Furthermore, the measurement campaigns selected took place in remote areas with surface ozone concentrations around 30–40 ppb ozone and NO$_2$ typically well below 1 ppb. The OD associated with NO$_2$ was $(2 \pm 9) \times 10^{-4}$, ozone below $1.5 \times 10^{-4}$ and iodine monoxide $(2.2 \pm 2.2) \times 10^{-4}$. In comparison, the typical OD of water vapour for the wavelength intervals W0–W5 ranged up to $1.65 \times 10^{-2}$ (see Fig. 1). Since the measurements were performed on the ocean, the influence of possible VRS in liquid water on the MAX-DOAS results was tested for and not found to be significant. Changes in water vapour dSCDs were about 1% when including liquid water Raman scattering in the fit.

For the LP-DOAS measurements the same upper limits for the absorption of NO$_2$ were found. The detection limits for O$_3$, glyoxal and IO were 191 ppb, 30 and 0.36 ppt, which were not exceeded during our nightly measurements. Therefore these trace gases were not included in the final analysis.

4.4.2 Saturation effects

The saturation effect originates from the fact that the convolution of literature cross-sections with the instrument function does not commute with the exponential function in the Lambert–Beer Law (Wenig et al., 2005). This effect can be corrected for by replacing the absorption cross-section $\sigma(\lambda)$ by $\sigma_{\text{Sat}}(\lambda)$ given by Eq. (3) for a given column density SCD.

$$\sigma_{\text{Sat}}(\lambda) = \frac{1}{S} \ln(\exp(-S \times \sigma(\lambda)) \otimes H)$$

The saturation effect for a dSCD of $6 \times 10^{23}$ molec cm$^{-2}$ would reduce the apparent OD for the strongest absorption discussed here at 442 nm by about 2%. A significant influence of the saturation effect on the results presented here can therefore be ruled out. This consideration is supported by the observed linear relationship between the retrieved water vapour SCDs retrieved for weaker (W1) and stronger (W3) absorption bands shown in Fig. 3.

4.4.3 The solar $I_0$ effect

The solar $I_0$ effect (Platt et al., 1997) describes the effective weighting of the absorption cross-section when the spectrum of the light source is highly structured itself, such as that of the Sun.

This needs to be corrected for MAX-DOAS measurements only, since the light source of LP-DOAS systems have usually broader structures than the absorbers itself. The zeroth-order $I_0$ correction does therefore not rely on the OD of the absorber but can be corrected together with the saturation effect as described by Aliwell et al. (2002) and Vogel et al. (2013) for a fixed dSCD. For our measurements corrections were made by calculating a modified $\sigma^*_0(\lambda)$ using the Kitt Peak solar flux atlas (Chance and Kurucz, 2010):

$$\sigma^*_0(\lambda) = \frac{1}{S} \ln \left( \frac{(I_0 \exp(-S \times \sigma(\lambda))) \otimes H}{I_0 \otimes H} \right).$$

The $I_0$ effect can result in changes of apparent cross-section in MAX-DOAS measurements of up to 10%. Because these changes can reduce as well as enhance the apparent OD, the data were evaluated with and without solar $I_0$ correction to estimate its effect on the relative absorption strength of different water vapour absorption bands. The data shown in Table 5 are $I_0$-corrected and the $I_0$ effect accounts here for changes of at most 2%. Including the $I_0$ effect resulted in a significant reduction of the residual and therefore the fit errors. For the MAX-DOAS evaluation, a SCD for the $I_0$ correction of $1 \times 10^{15}$ molec cm$^{-2}$ for NO$_2$, $1 \times 10^{18}$ molec cm$^{-2}$ for O$_3$ and $3 \times 10^{23}$ molec cm$^{-2}$ for H$_2$O was applied.

5 Discussion

From the fit errors listed in Table 6 especially for the weak absorption features of water vapour, it can be seen that the development of water vapour absorption compilations from HITRAN 2009 to HITRAN/HITRAN 2012 results in a better fit of the measurement data. The fit errors for the intervals (W0 to W5) of the water vapour absorption cross-section are reduced. Changing from HITRAN 2009 to HITEMP or BT2 reduces nearly all fit errors somewhat; the reduction is dramatic (20–75%) for the weaker bands (W1, W2 and W4). The smallest relative fit errors are observed for most absorption bands for HITEMP.

The magnitude of the main absorption W3 at 442 nm agrees with data from a meteorological station as shown in Table 8.

Nevertheless, the relative absorptions of different groups of absorption lines are inconsistent and do not fit our measurements; they are listed relative to W3 in Table 5. In the BT2 line list and in HITEMP the absorptions from 410 to 434 nm overestimate the observed absorptions approximately by a factor of 2.

Tennyson (2014) noted that the absorption line intensities for transitions involving highly excited vibrational states can depend strongly on the representation of the dipole moment surfaces in the ab initio models, particularly for transitions involving bending modes. This could be in agreement with our observations, since the largest discrepancy between modelled and measured absorptions was observed for the $7\nu + \delta$ polyad.
As it turned out in Sect. 4.1 a separation of the cross-section into separate cross-sections for different bend mode values \( v_2 \) indicates that the difference in relative absorption strength can be attributed to absorption lines with higher bend mode values. Their intensities are systematically overestimated (Table 7).

- For W0 (394–410 nm) the results from MAX-DOAS and LP-DOAS agree that the magnitude of the absorption at W0 is found to be about 5 ± 6% smaller in measurements than what is reported in HITEMP.

- For W1 (410–423.5 nm) the agreement of MAX-DOAS and LP-DOAS measurements is not as good, since the overall absorption is about half as large as W0. The absorption of this group of absorption lines is too high by 60 ± 10% in HITEMP. For W1 in BT2 a better agreement is found, together with smaller fit errors. Even when considering the overestimation of the absorption W1 in HITEMP, the fit error was reduced from HITRAN 2009 to HITEMP. This shows that the shape of the absorption is reproduced more accurately in HITEMP.

- The absorption W2 (423.5–434 nm) was not regularly identified in LP-DOAS measurements and is also close to the magnitude of the residuals for MAX-DOAS data. Therefore the result for W2 should be viewed with caution. The shape of the water cross-section is reproduced correctly within 1 × 10^{-27} cm^2 at 0.45 nm resolution. HITEMP overestimates this absorption compared to observations by 30 ± 25%, BT2 by even 82 ± 33%, but its shape is reproduced better in HITEMP than in HITRAN 2009, as seen from the smaller fit errors.

- The absorption W3 (434–451.5 nm) is relatively strong and therefore requires application of saturation and \( I_0 \) corrections. Furthermore, neglecting the changes in radiative transfer for MAX-DOAS measurements for individual absorption lines leads to significant structures in the residual spectra (up to 5 × 10^{-4}), while the overall water vapour dSCD is only slightly changed by <3% for \( S < 5 \times 10^{23} \) molec cm^{-2}. The simple approximation of having a constant light path for the MAX-DOAS observation within this wavelength interval does not hold here for larger OD; therefore the fit errors also do not decrease, as seen for the BT2 list for LP-DOAS measurements in Table 6.

- The absorption W4 (451.5–461.5 nm) is small, but absorptions in this wavelength range could interfere with measurements of glyoxal. In fact the optical density due to H2O is comparable to previous observations of glyoxal on the open ocean (e.g. Mahajan et al., 2014): dSCD_{glyoxal} = 1 × 10^{13} molec cm^{-2} corresponds to 6 × 10^{-4} at 455 nm, dSCD_{H2O} = 4 × 10^{23} molec cm^{-2} to 4 × 10^{-4} for HITEMP at 456 nm) which also absorbs in this spectral region. A water vapour dSCD-correlated structure in the residuals is found at the absorption at 453.0 nm (HITEMP), which can also be seen in Fig. 2. This absorption seems to be better reproduced in BT2, the absorption is at 452.5 nm. The same error estimate for the convolved cross-section as for W2 applies for W4.

- For W5 (461.5–480 nm) good agreement for water vapour absorption based on HITEMP, HITRAN 2009 and BT2 was found in observed LP-DOAS spectra within an error of 4%. W5 was not within the measured wavelength range of the MAX-DOAS instrument.

On the basis of the observed discrepancies in relative absorption band strength, we suggest rescaling the respective water vapour absorption cross-section or including only wavelength intervals in a DOAS analysis where the relative absorption band strengths are sufficiently in agreement with each other. This means e.g. for IO that the water absorption band at 442 nm (W3) should be avoided when the absorption at 426 nm (W2) or even also at 416 nm (W1) is included. For the retrieval of glyoxal with its main spectral absorption features above 440 nm, a wavelength window which does not include water vapour absorption at 426 and 416 nm should be preferred when using these water vapour absorption cross-sections. The same argument applies for choosing a retrieval interval for NO2.

The relative strength of the absorption band W5 around 470 nm and the small fit errors indicate that the water vapour absorption cross-section is unlikely to cause interferences when retrieving O4 dSCDs for radiative transfer modelling in this spectral region.

The water absorptions included in HITEMP below 390 nm were observed in our measurements and found to be smaller than \( \sigma_{\text{max}} = 4 \times 10^{-27} \) cm^{2} for the 377 nm region for a spectral resolution of 0.45 nm. Absorptions observed in the 362 nm region might be caused by water vapour or by uncertainties of \( \sigma_{\text{O4}} \). Here absorptions with ODs of \( 4 \times 10^{-4} \) correspond to water vapour dSCDs of \( 3 \times 10^{23} \) molec cm^{-2} (BT2) or \( 6 \times 10^{23} \) molec cm^{-2} (HITEMP), which are realistic dSCDs in mid-latitude regions. Typical \( \sigma_{\text{O4}} \) absorptions in MAX-DOAS measurements in this regions are of the order of \( 2 \times 10^{-2} \). Since the \( \sigma_{\text{O4}} \) absorption is used to retrieve aerosol extinction profiles from MAX-DOAS data, it might be necessary to correctly compensate for water vapour absorption to obtain correct profile information. This would require more theoretical calculations and dedicated validation measurements of the water vapour absorption cross-section below 395 nm.

Another aspect which needs to be considered is the wavelength dependence of the air mass factor in MAX-DOAS measurements covering several significant absorption bands, which implies that low residuals can only be obtained when accounting for these effects. For example, for a water vapour...
dSCD of $5 \times 10^{23}$ molec cm$^{-2}$ these effects will lead to residual structures of $\approx 2.5 \times 10^{-4}$ when including W1–W3 or already $\approx 1 \times 10^{-4}$ within W3 itself (compare also Table 4).

### 6 Conclusions

The revised line compilations HITEMP and BT2 lead to considerable improvements in DOAS measurements of water vapour and other trace gases. In particular, lowering the threshold above which absorption lines are included in the HITRAN database as well as theoretical and experimental progress led to improved water vapour absorption cross-sections in the blue wavelength range.

We showed that the water vapour absorption fit errors of the DOAS retrieval can be reduced by up to 20–75% by using current water vapour absorption cross-sections (e.g. HITEMP: Rothman et al., 2010) compared to older versions of HITRAN. This shows that the shape of the water vapour absorption is better reproduced by the HITEMP, reducing the amplitude of residual spectra. When the error in the relative magnitude of several absorption bands are corrected (e.g. by scaling the water vapour absorption cross-sections in the different spectral intervals), the amplitude of residual spectra for larger fit intervals can be reduced significantly (compare e.g. Fig. 2), which then allows more reliable retrieval of weak absorbers.

Each of the water vapour absorption cross-sections evaluated here (HITRAN 2009, HITRAN 2012, HITEMP and BT2) was found to have certain limitations for the use in DOAS retrievals in the blue wavelength range. Even after water absorption cross-sections have improved they are still not sufficiently accurate for modern high-precision DOAS measurements of tropospheric trace gases at mid-latitudes. Fit errors obtained from evaluating measurement data were found to become lower for individual absorption features with the improvement of the HITRAN database from HITRAN 2009 to HITEMP and HITRAN 2012. However, we also showed that there are still problems concerning the relative strength of the different absorption bands in the blue wavelength range which need to be addressed.

The correction factor for different absorption bands from 394 to 480 nm, which we derived from our atmospheric measurements listed in Table 9. They range from 0.5 to 1.9 and are highly significant. Even the most recent water vapour absorption cross-section (HITRAN 2012) still requires the application of correction factors ranging between 0.63 and 1.0. The corrections are necessary for all high-precision DOAS measurements in this spectral range when significant water vapour absorption is present. The here-presented correction factors can be used until better absorption line lists are available. Inclusion of even weaker absorption lines in the databases could further improve the modelling of water vapour optical densities, as indicated by the minimum in fit errors of LP-DOAS data when using the BT2 line list without a threshold for the absolute magnitude of the cross-section of individual absorption lines. From our measurements, water vapour absorptions below 385 nm remain uncertain and do not match current water vapour absorption cross-section data. We could not confirm recent UV water vapour absorption cross-section measurements by Du et al. (2013) and found upper limits for the absorption cross-section which are 2 orders of magnitude lower. Therefore further research to provide more consistent water vapour absorption cross-section data in the UV/VIS range is necessary.

High-quality LP-DOAS measurements along light-path lengths of $\approx 10$ km and a comparably high spectral resolution in a tropical climate seem to be feasible in order to investigate further the relative absorption band strength without the need to correct for radiative transfer effects, especially the study of different vibrational states outlined in Sect. 4.1.

**Acknowledgements.** We thank J. Tennyson, H. Sihler and K. Großmann, R. Volkamer, T. Koenig, the editor and the two referees for helpful comments during the preparation of the manuscript.

---

**Table 9.** Correction factors for HITRAN 2009, HITEMP and BT2, calculated as error-weighted means of the values derived from MAX-DOAS and LP-DOAS observations listed in Table 5 relative to W3. Values listed in square brackets in Table 5 were not used. When both values in Table 5 are above a relative fit error of 25%, only the MAX-DOAS values are used and listed in square brackets. The relative error for each of these values listed in Table 5 was estimated to be twice the fit error listed in Table 6.

<table>
<thead>
<tr>
<th>Dominating polyad</th>
<th>$8\nu$</th>
<th>$7\nu + \delta$</th>
<th>$7\nu$</th>
<th>$6\nu + \delta$</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>W0</td>
<td>W1</td>
<td>W2</td>
<td>W3</td>
</tr>
<tr>
<td>Start of interval (nm)</td>
<td>394.0</td>
<td>410.0</td>
<td>423.5</td>
<td>434.0</td>
</tr>
<tr>
<td>End of interval (nm)</td>
<td>410.0</td>
<td>423.5</td>
<td>434.0</td>
<td>451.5</td>
</tr>
</tbody>
</table>

<table>
<thead>
<tr>
<th>Source of cross-section data</th>
<th>$\delta$</th>
<th>$\pm$</th>
<th>$W_0$</th>
<th>$W_1$</th>
<th>$W_2$</th>
<th>$W_3$</th>
<th>$W_4$</th>
<th>$W_5$</th>
</tr>
</thead>
<tbody>
<tr>
<td>HITRAN 2009</td>
<td>1.06</td>
<td>0.07</td>
<td>1.95</td>
<td>0.04</td>
<td>[1.69]</td>
<td>1.14</td>
<td>1.02</td>
<td>0.07</td>
</tr>
<tr>
<td>HITEMP</td>
<td>1.05</td>
<td>0.06</td>
<td>0.63</td>
<td>0.04</td>
<td>0.77</td>
<td>0.15</td>
<td>1.03</td>
<td>0.07</td>
</tr>
<tr>
<td>BT2</td>
<td>1.06</td>
<td>0.05</td>
<td>0.48</td>
<td>0.03</td>
<td>0.55</td>
<td>0.10</td>
<td>1.01</td>
<td>0.06</td>
</tr>
</tbody>
</table>

---

**Atmos. Meas. Tech., 8, 4329–4346, 2015**
We thank the captain, officers and crew of *Meteor* for support during cruise M91. We thank the CVAO team, especially L. Mendes, for support during the HALOCAVE campaign. We thank GEOMAR for logistical support. We thank the German Science foundation DFG within the core program METEOR/MERIAN. We thank the German ministry of education and research (BMBF) for supporting this work within the SOPRAN (Surface Ocean Processes in the Anthropocene) project (Förderkennzahl: FKZ 03F0662F) which is embedded in SOLAS. We thank the University of York/NCAS (National Centre for Atmospheric Science) for providing meteorological measurement data at CVAO. We thank the authorities in Peru for the permission to work in their territorial waters.

The article processing charges for this open-access publication were covered by the Max Planck Society.

Edited by: A. Richter

### References


Shefov, N. N.: Intensivnosti nokotorykh emissiy sumerochnogo i nochnogo neba (Intensities of some Emissions of the Twilight and Night Sky), Spectral, electrophotometrical and radar researches of aurorae and airglow, IGY program, section IV, 1, 25, 1959.


